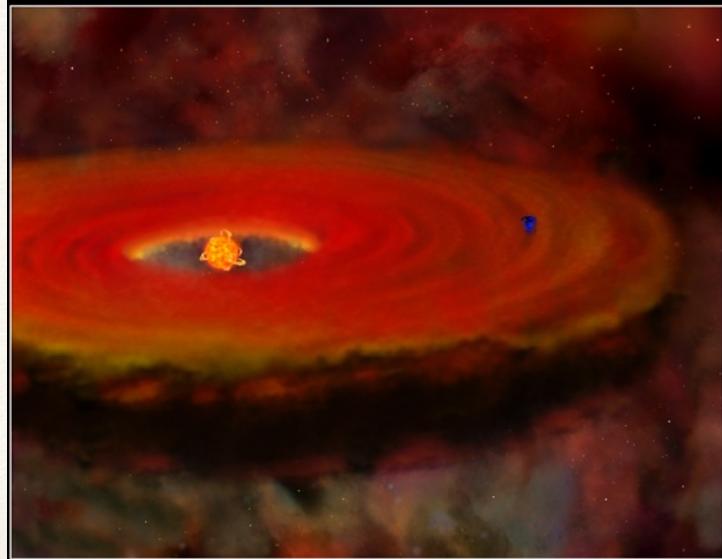
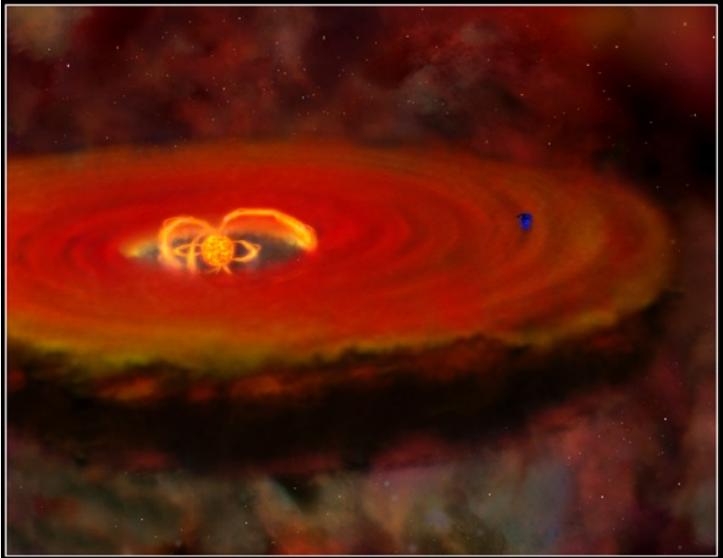


(c) NASA/CXC/M. Weiss



# The end of an era? A search for flickering accretion in T Tauri stars

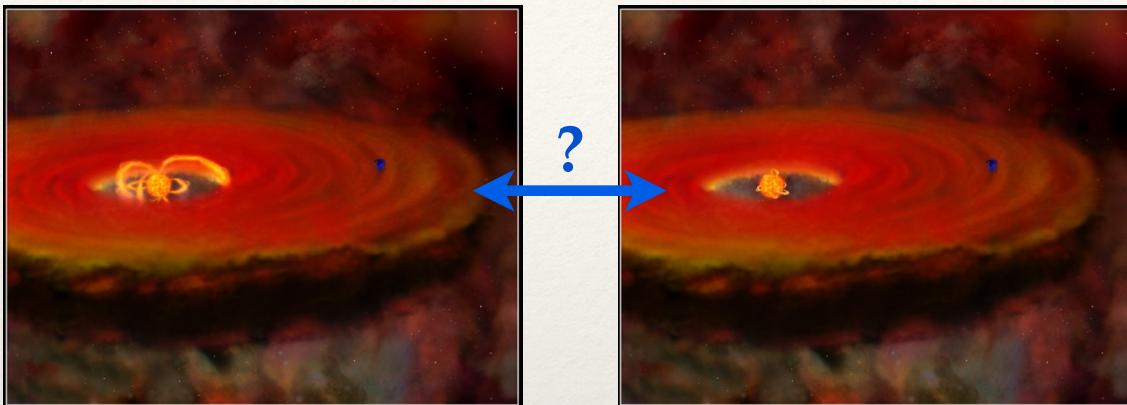
Gaspard Duchêne (UCB, IPAG)



*Adam Becker, Tony Yang (UCB),  
Hervé Bouy (CAB Madrid), Jenny Patience (ASU),  
Rob de Rosa (UCB)*



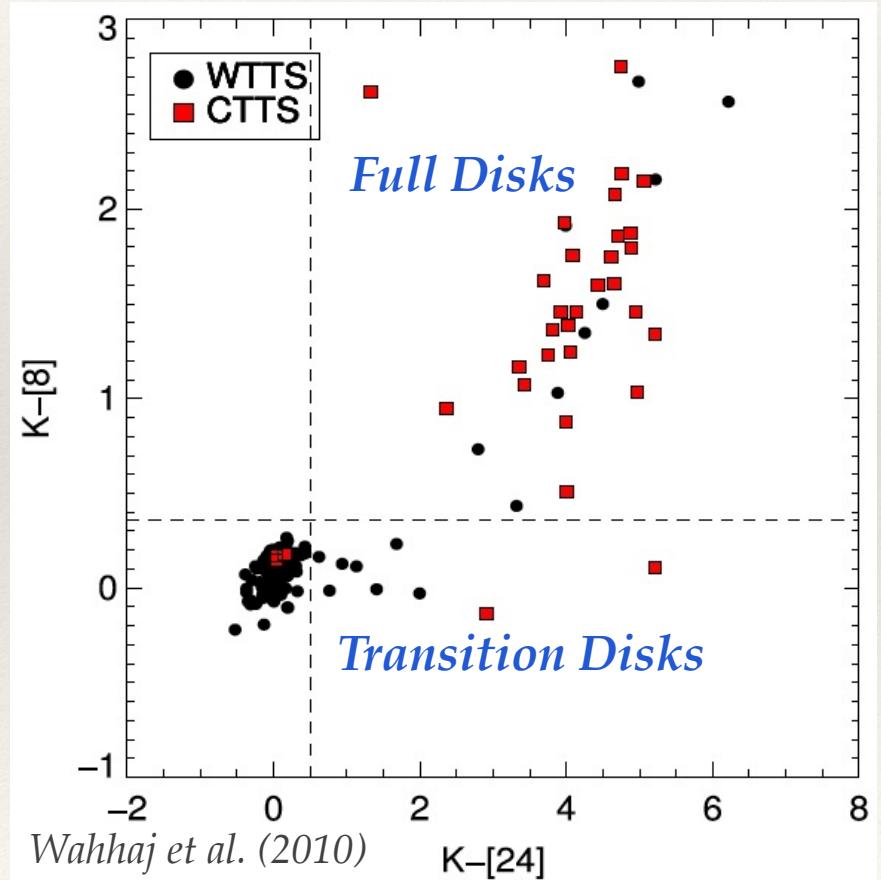
# Protoplanetary disks & Accretion



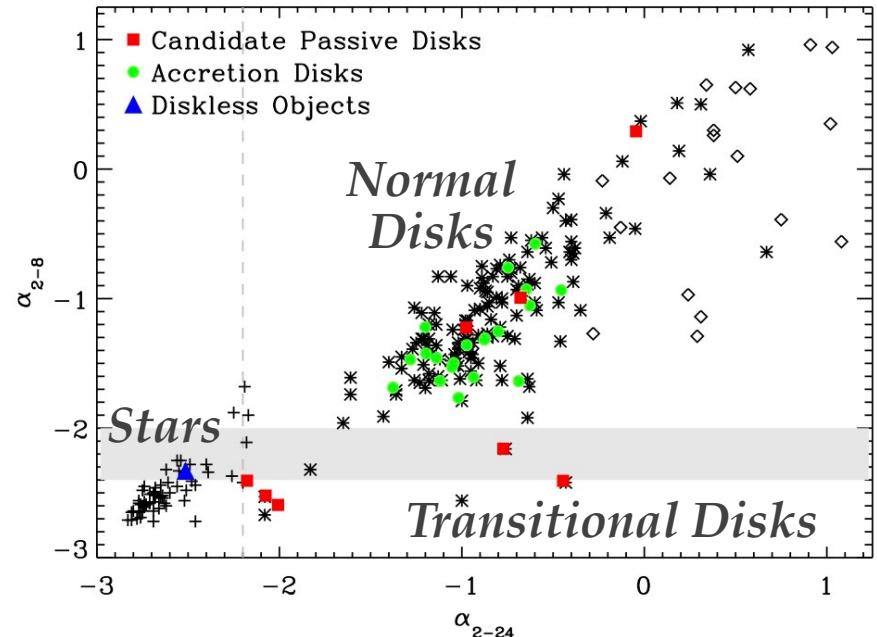
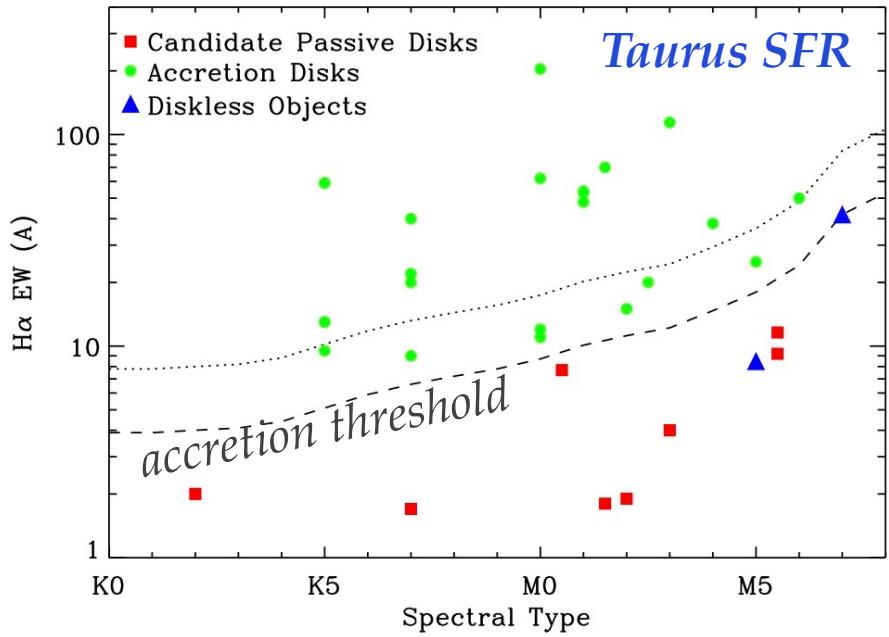
- Misclassified Weak-Line TTS?
- Intermittent accretion?
- Purely passive disks?
- **Initial stages of disk dissipation?**

*Optical spectroscopic monitoring survey of accreting and apparently non-accreting T Tauri stars*

Some ( $\approx 10\%$ ) dusty disks show no accretion! Why??



# Sample, set-up and strategy

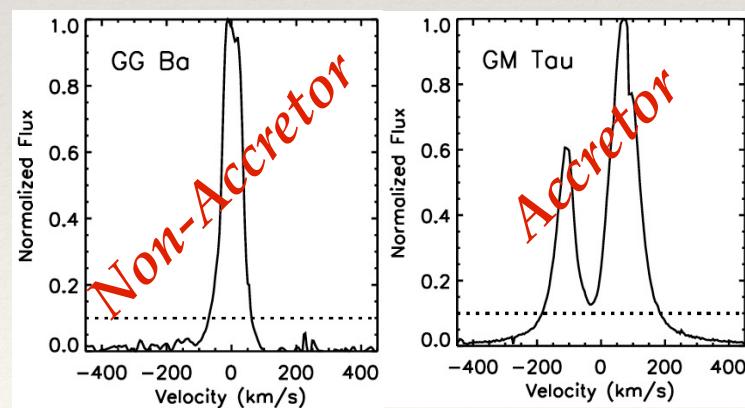


## Campaign details:

2-4m telescopes  
 R~2,000-5,000 spectra  
 8-12 epochs/target  
 $1 \text{ day} \leq \Delta t \leq 13 \text{ months}$

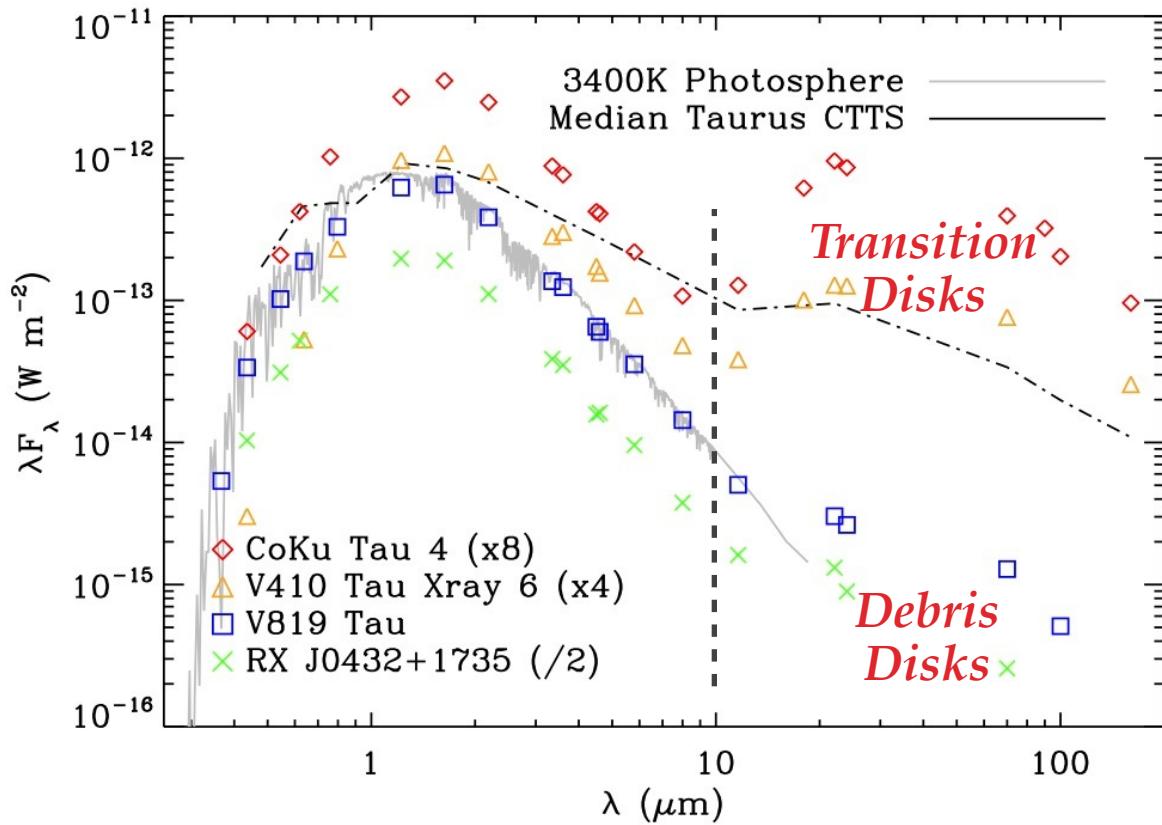
## Primary diagnostic:

H $\alpha$  10% width (W<sub>10%</sub>)

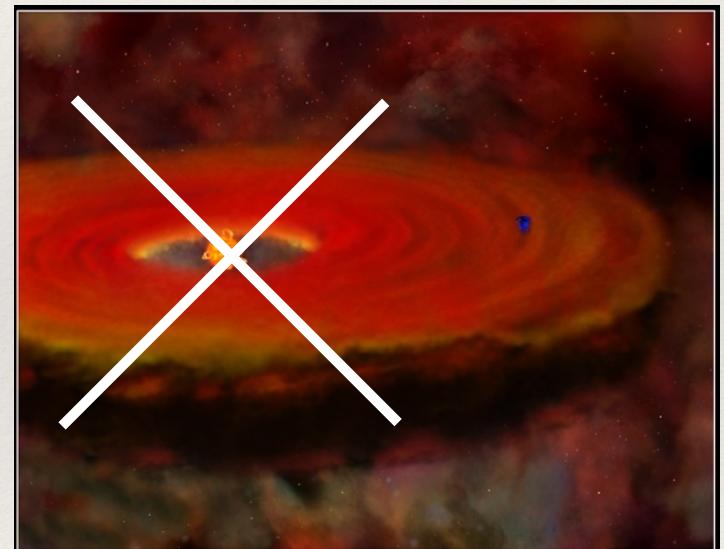


White & Basri (2003)

# No continuous passive disk!

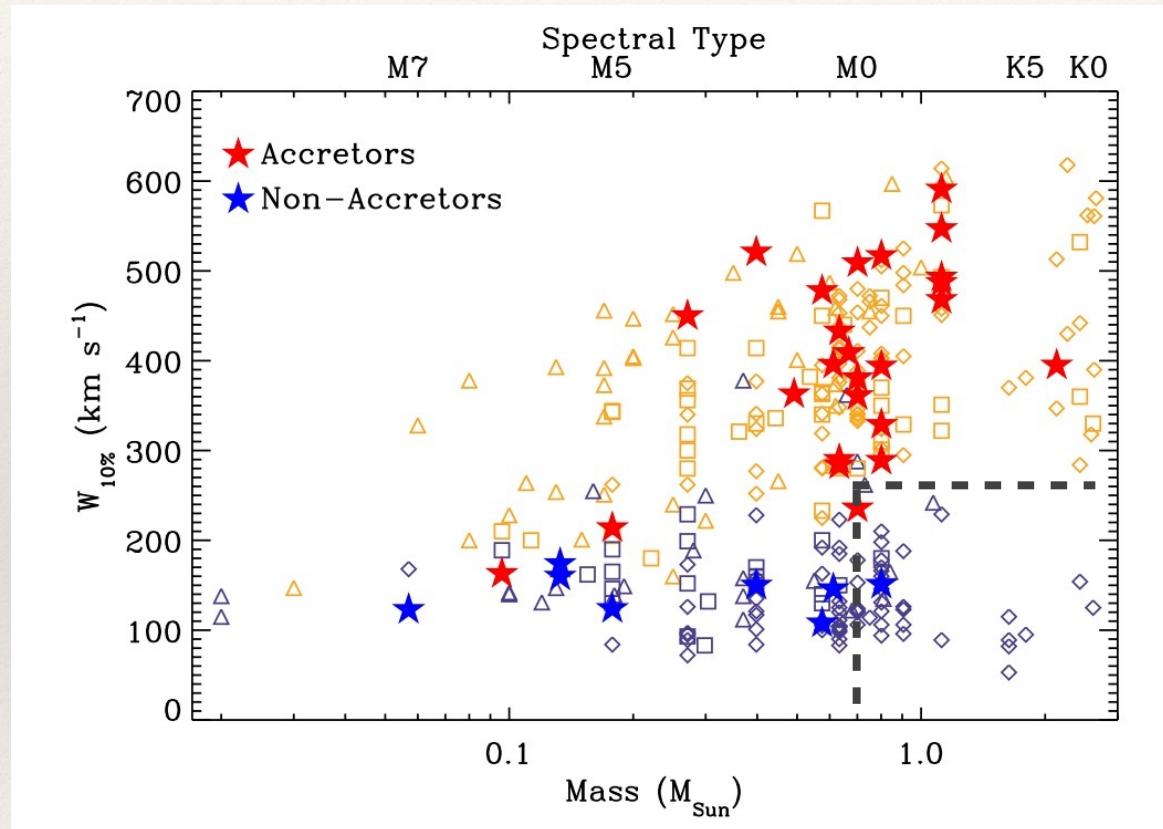


Fast inner disk dissipation  
of both dust and gas

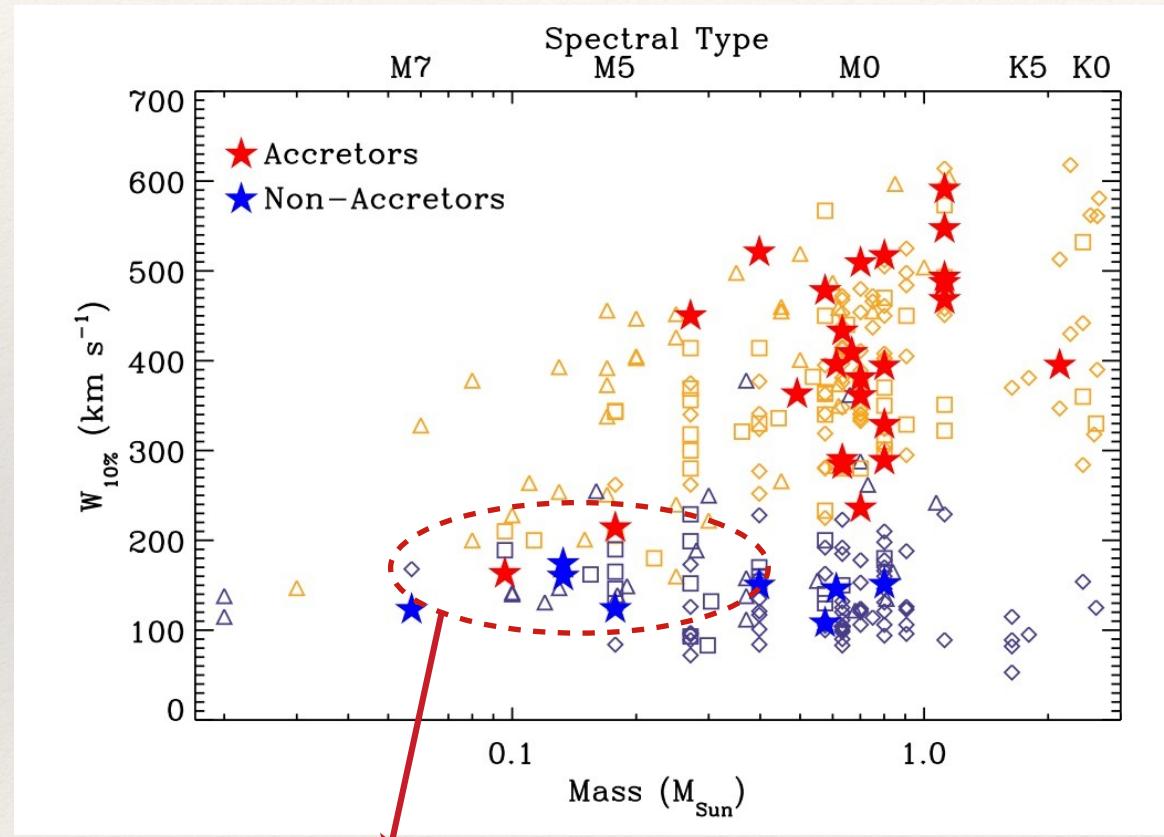


Several impostors: weak-but-broad-and-variable lines, multiple systems

# Does $W_{10\%}$ probe the accretion rate?

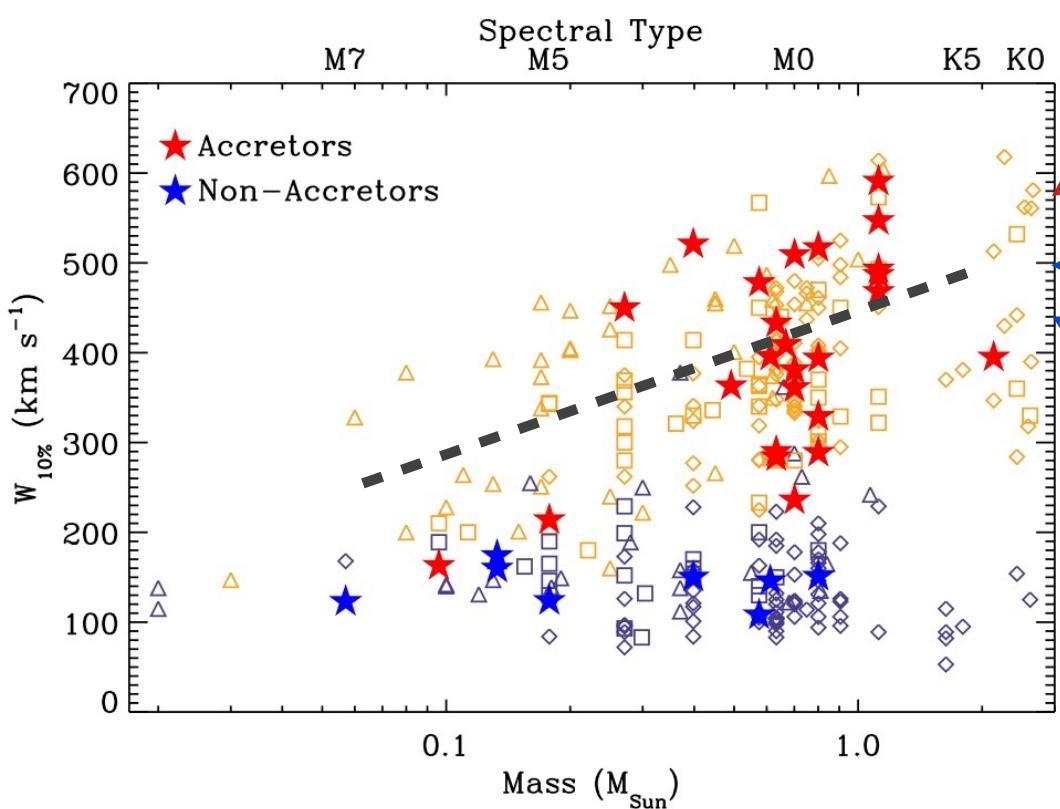


# Does $W_{10\%}$ probe the accretion rate?



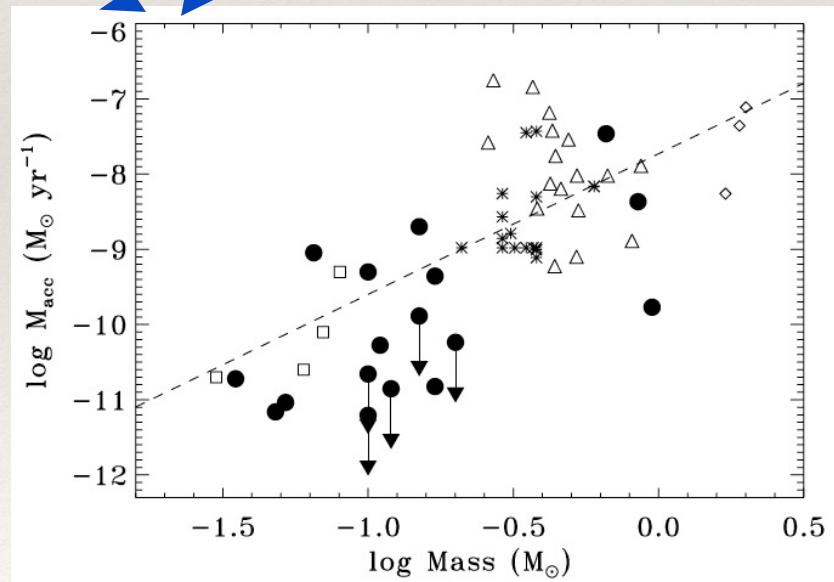
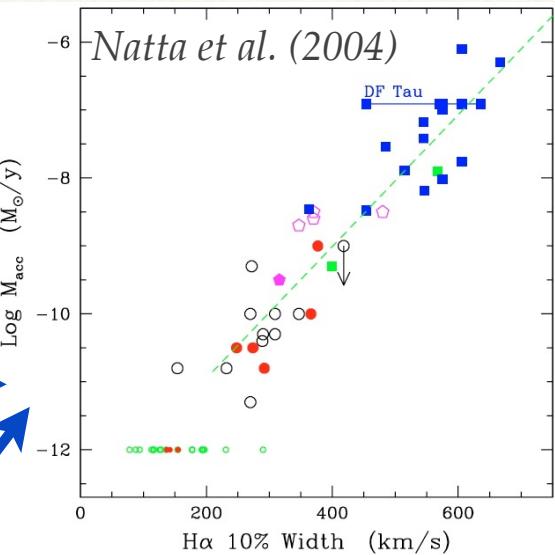
$W_{10\%}$  is an ambiguous criterion  
for low-mass TTS

# Does $W_{10\%}$ probe the accretion rate?



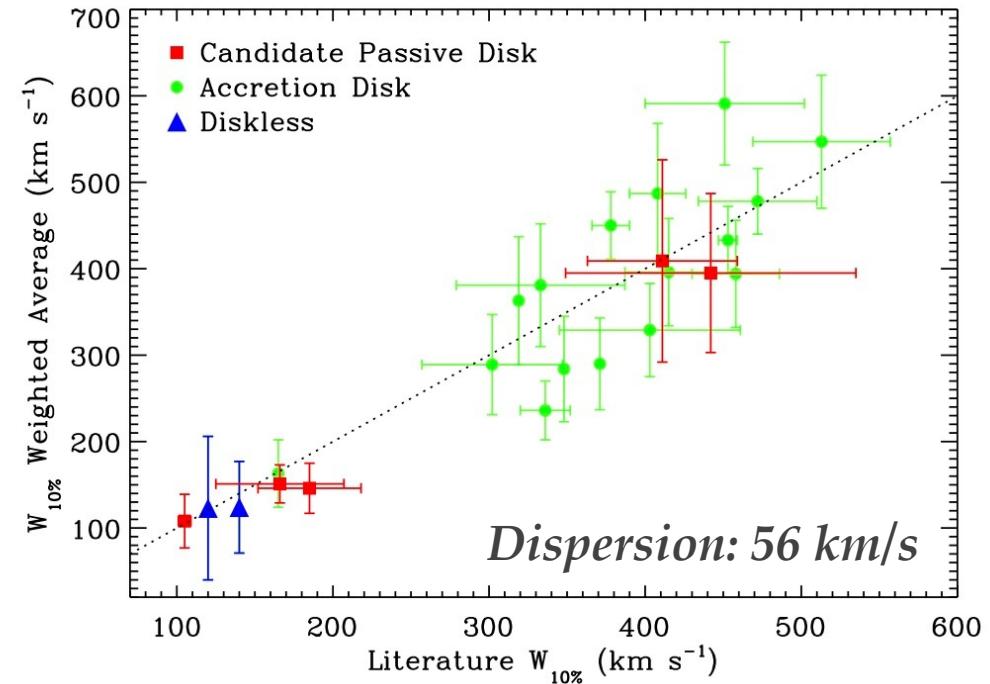
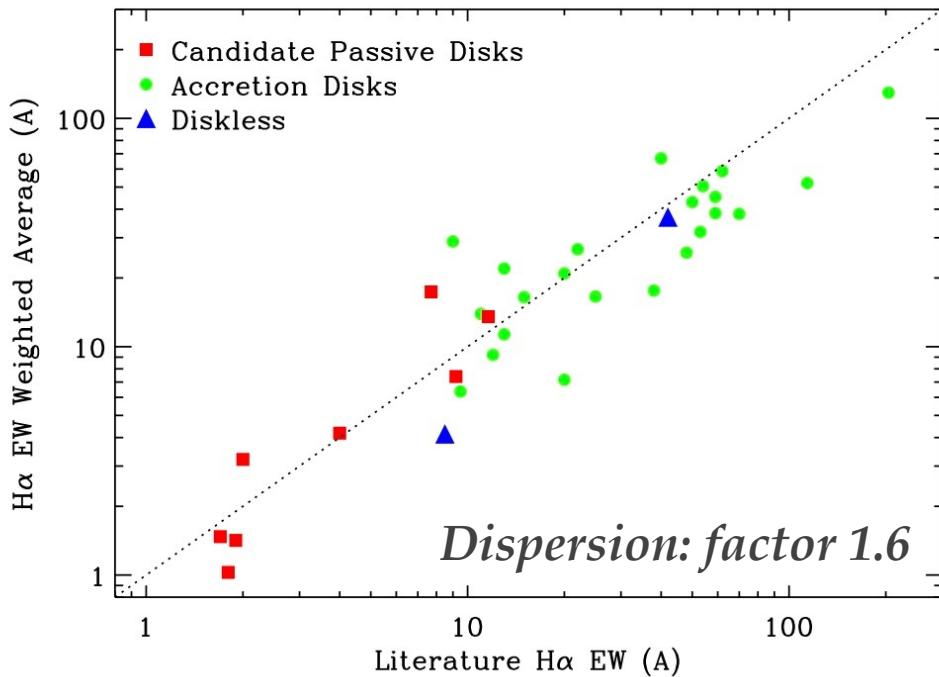
*Free-fall  
velocity?*

*Which one is  
physical?*



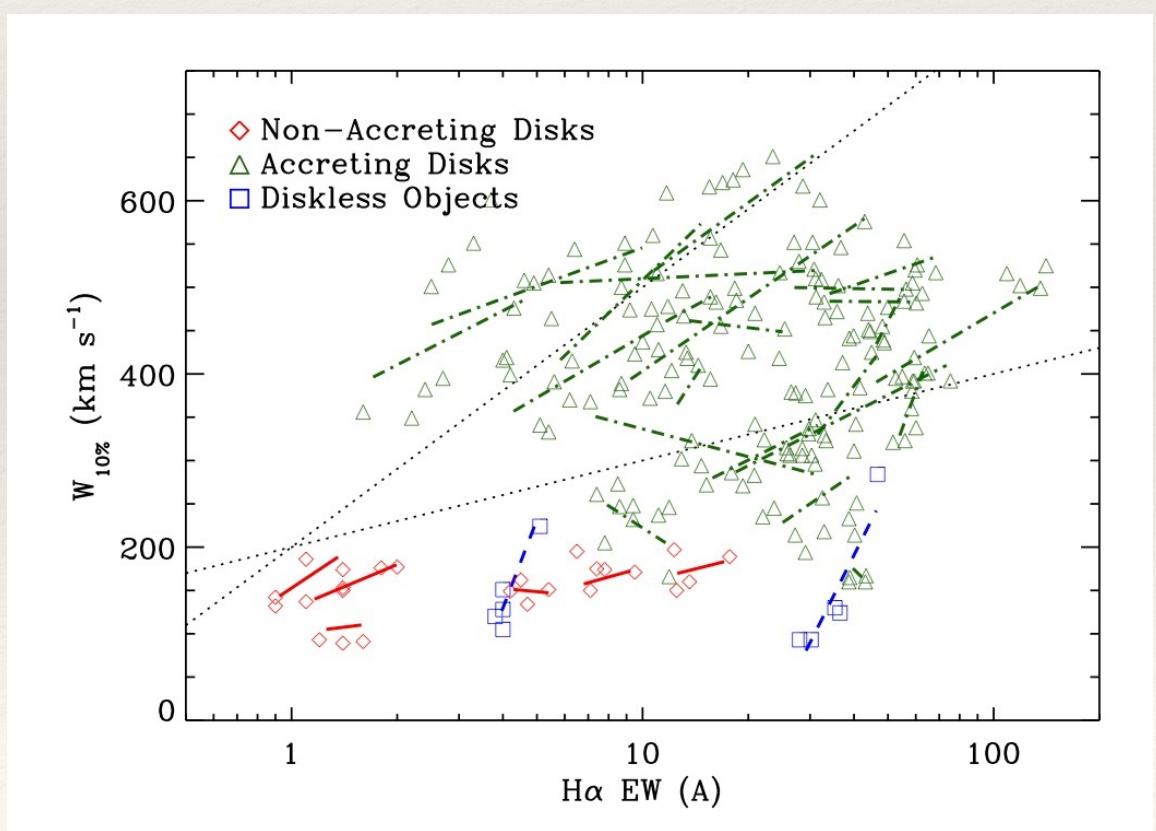
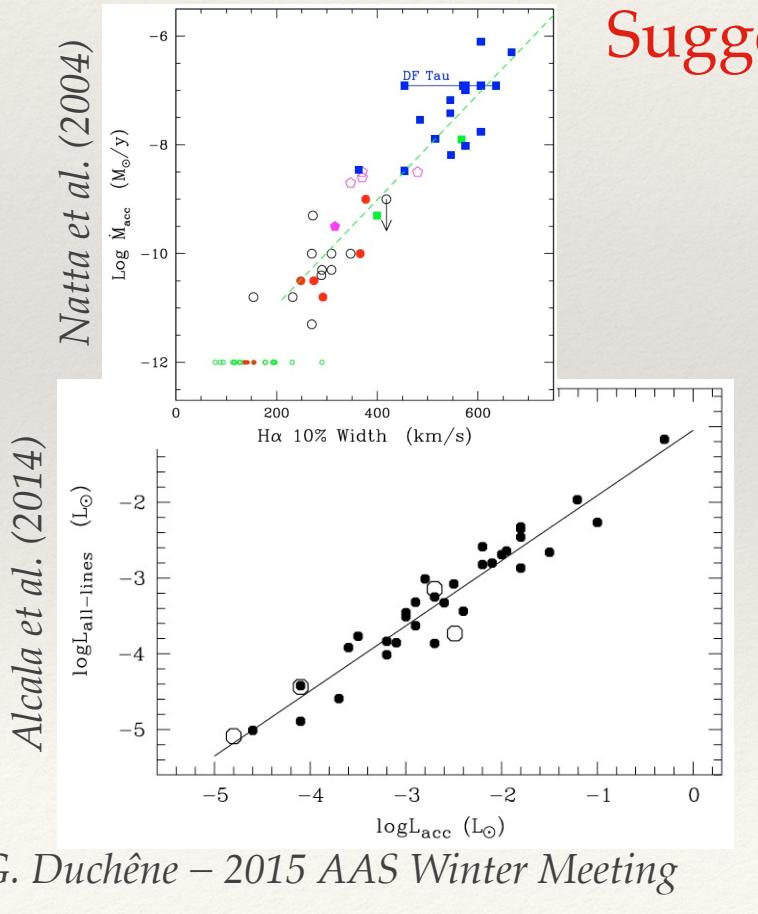
# Time variability: Secular trend

- ❖ Average H $\alpha$  line properties in line with previous obs.
- ❖ Suggests a constant *mean* accretion rate over several yr



# Time variability: Short timescales

- ❖ Excursions for most accreting objects follow a similar relationship as ensemble correlations predict  
*Suggests that  $W_{10\%} \propto \dot{M}_{\text{acc}}$  is physically driven*



# Conclusions

- ❖ Our 13-month monitoring spectroscopic survey of accreting and non-accreting T Tauri stars reveal:
  - no continuous passive disk
  - no object with flickering accretion
  - steady behavior of accretors over several yr
  - some physical grounding for  $W_{10\%} \propto \dot{M}_{\text{acc}}$
  - the ambiguity of  $W_{10\%}$  accretion threshold

