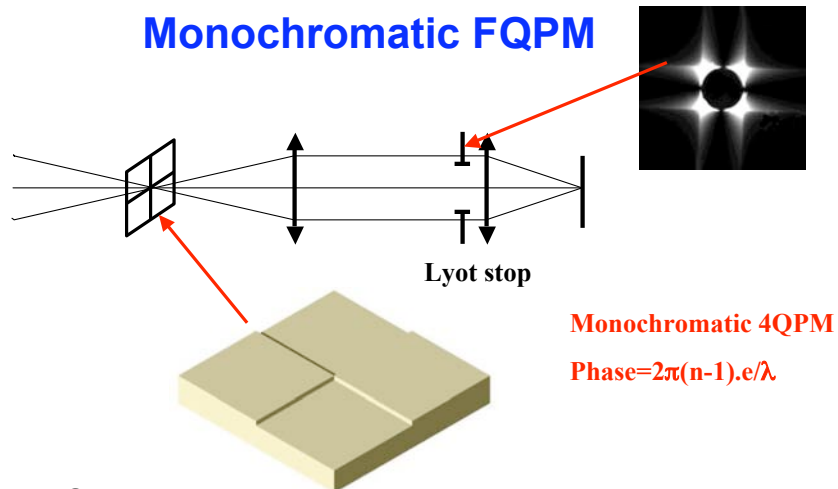


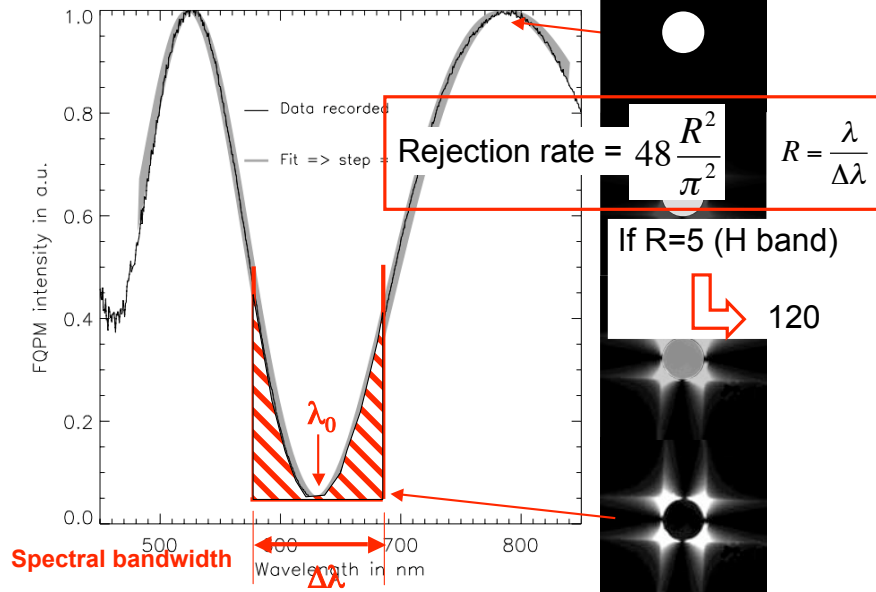
Multiple-Stage Four Quadrant Phase Mask Coronagraph

P. Baudoz, R. Galicher, A. Boccaletti,
D. Rouan, J. Baudrand

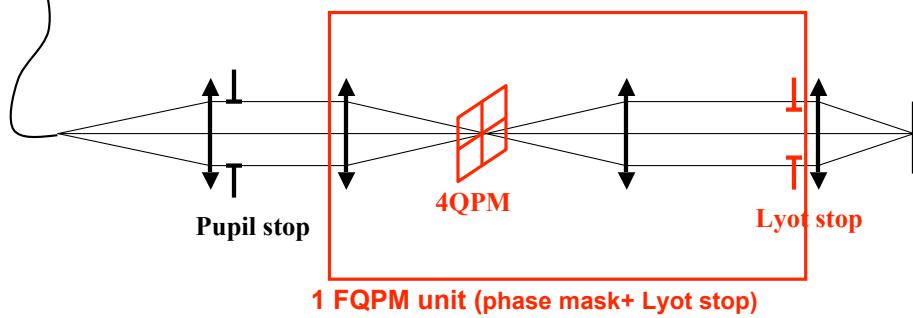


- Simple
- well studied (lab, JWST, VLT-NACO)
- already built for visible, NIR, thermal IR

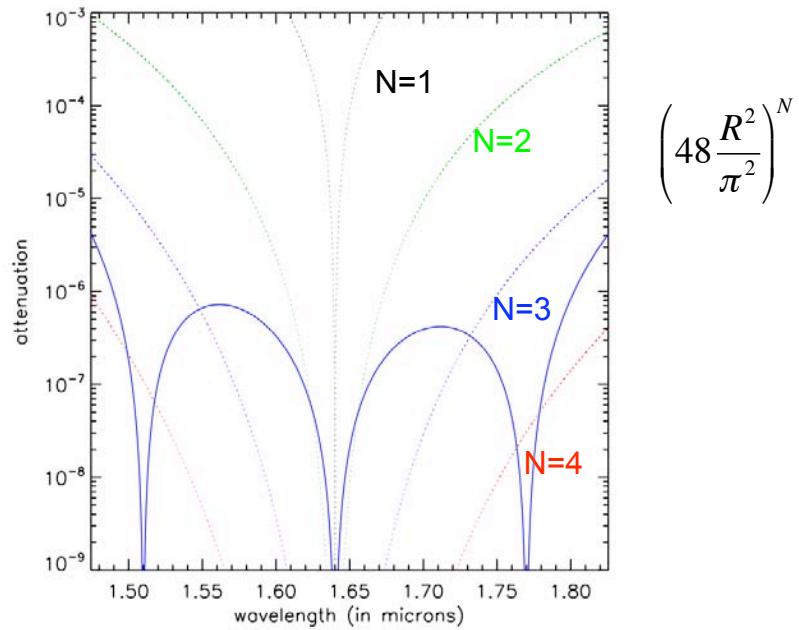
FQPM chromaticity



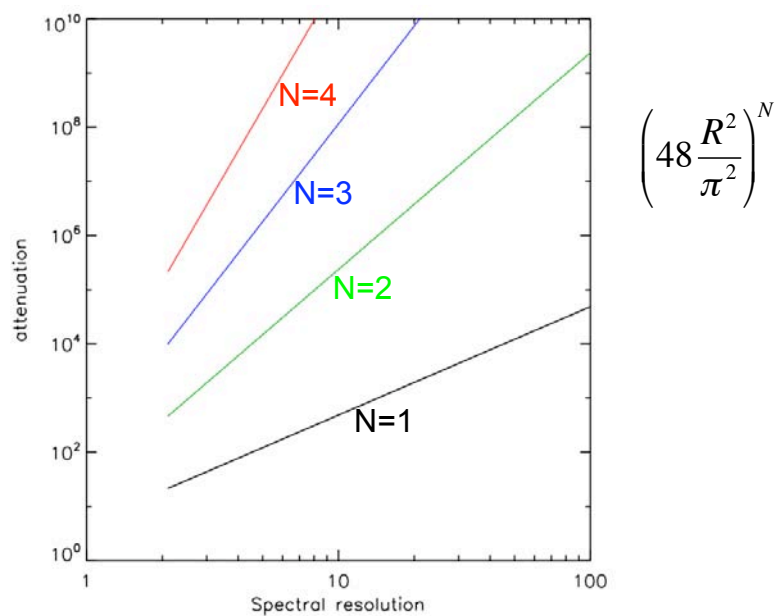
Principle of Multi-stage FQPM



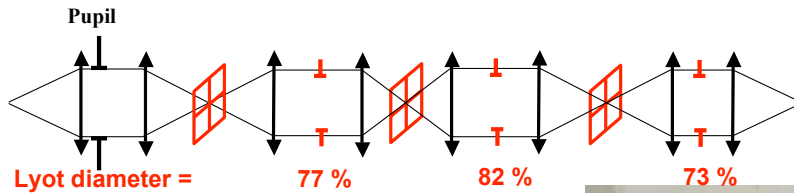
Multi-stage FQPM chromaticity (H Band)



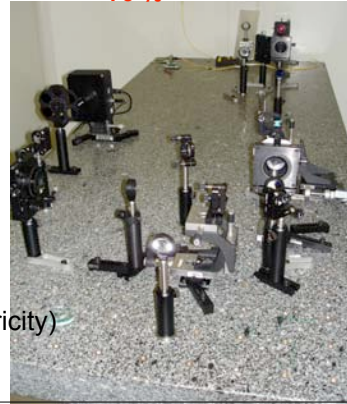
General Multi-stage FQPM chromaticity



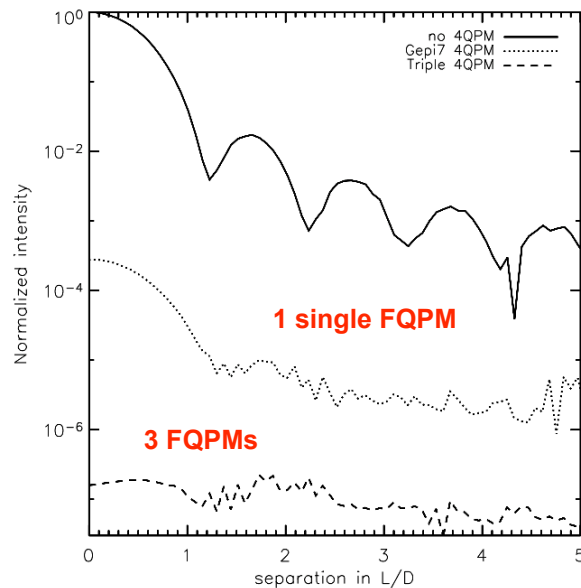
Laboratory tests : set-up



- Super-luminous source
- 3 monochromatic FQPMs optimized for:
 - 623 nm
 - 633 nm
 - 641 nm
- Standard lenses
- Entrance pupil : 1.78 mm (chromaticity, sphericity)

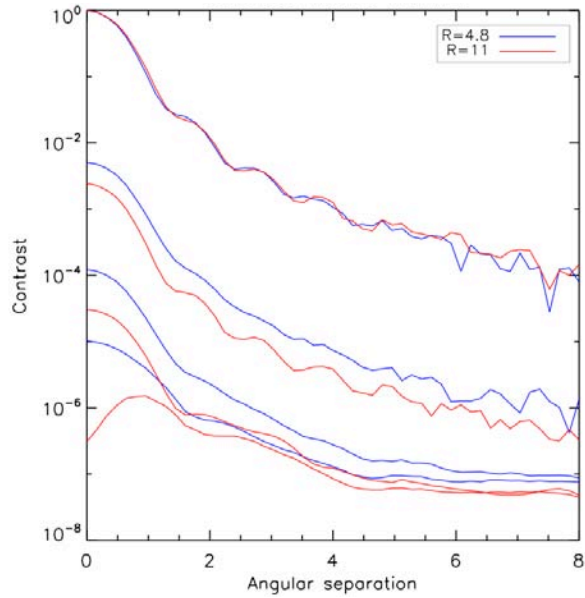


Laboratory tests :monochromatic

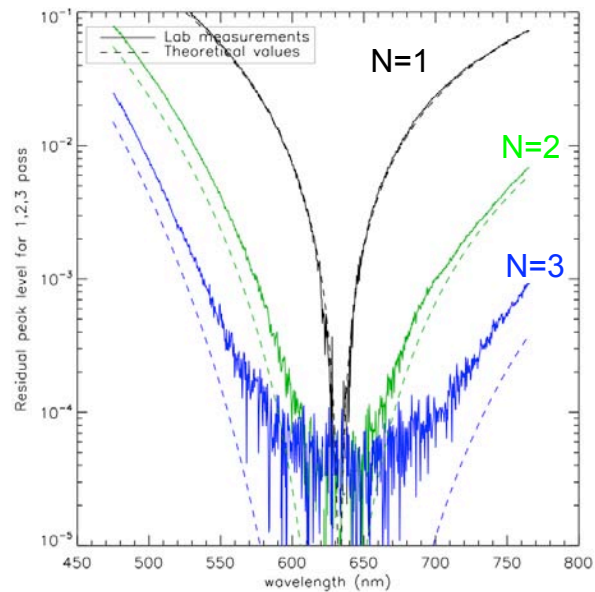


Profiles for Multiple Stage FQPMs

- N=1, 2 and 3
- Spectral bandwidth
 $R = \lambda / \Delta\lambda = 4.8$ and 11



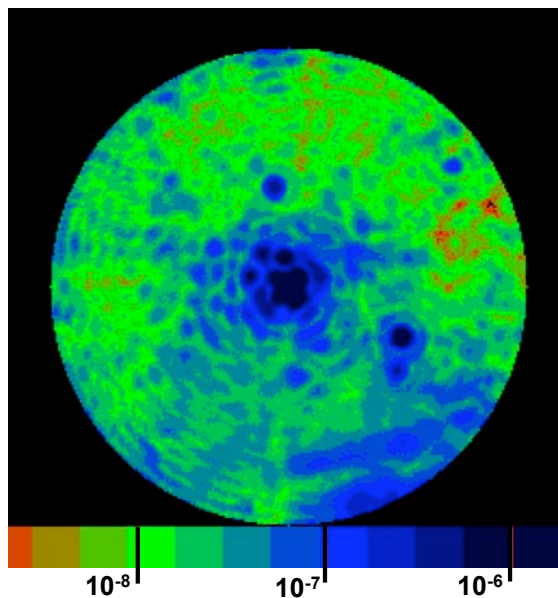
Laboratory tests : spectroscopy



Laboratory results $R=\lambda/\Delta\lambda=11$

3 FQPM

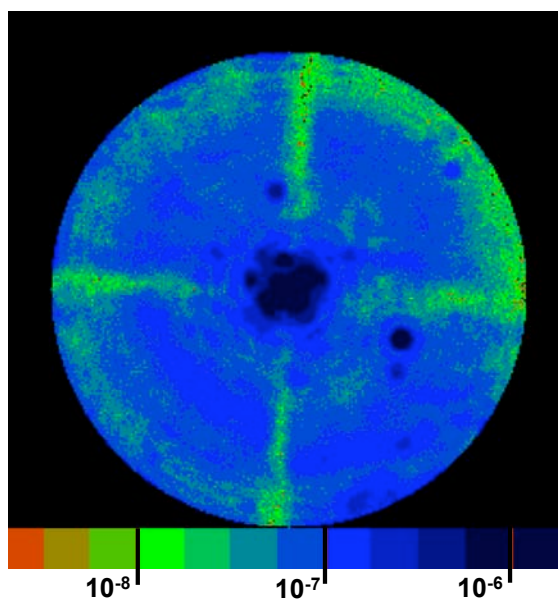
$35 \lambda/D$



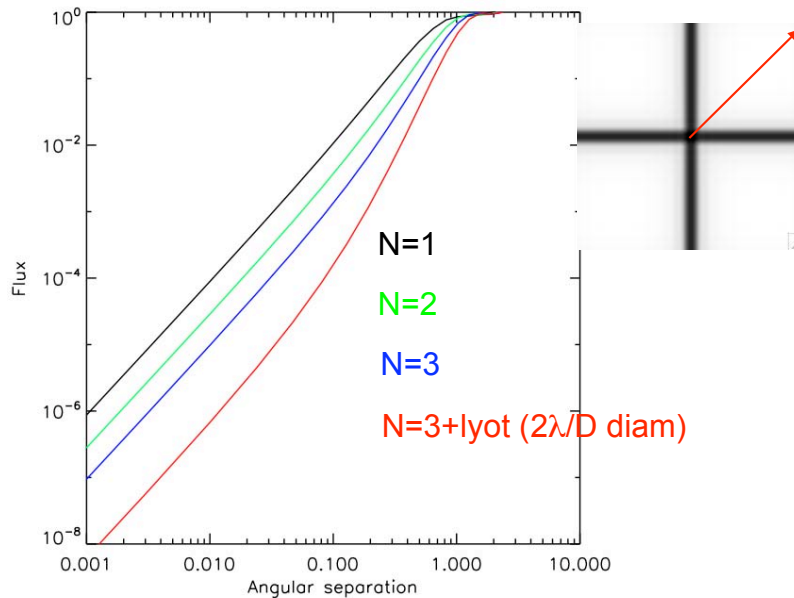
Laboratory results $R=\lambda/\Delta\lambda=4.8$

3 FQPM

$35 \lambda/D$

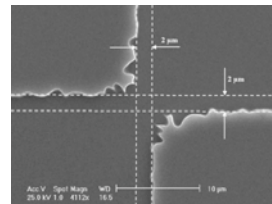


Multi-stage IWA



Advantages of Multiple-stage

- Based on well known technologies
- Small Inner Working Angle ($< 1.5 \lambda/D$)
- But relaxed stellar size issue
- Relaxed central obscuration issue
- Relaxed fabrication specification



Conclusion

Theoretically:

with 3 FQPMs well optimized $> 10^6$ rejection rate

Laboratory test :

10^7 @ $4 \lambda/D$ for typical astronomical bandwidth
 $R = \lambda/\Delta\lambda = 5$

To be done :

- full evaluation of performances with ground-based XAO
- design and build compact version with FQPM optimized for large spectral bandwidths.